

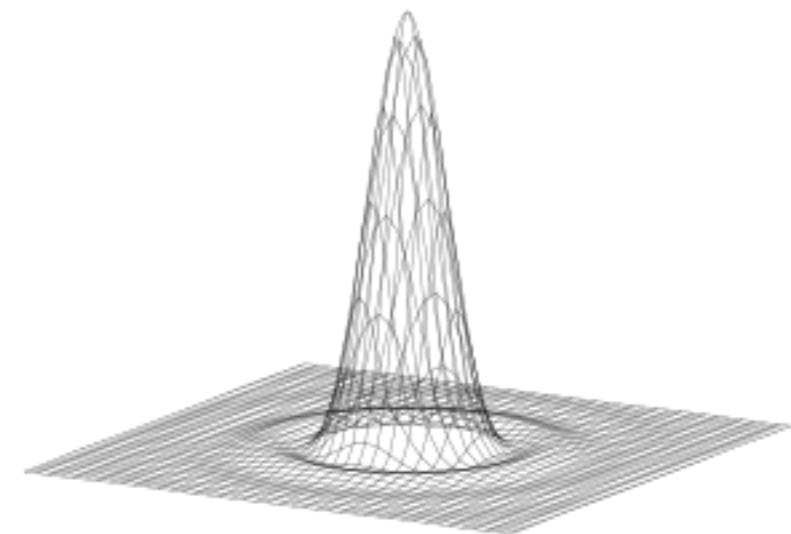


**Astrophysics Division  
CosmoStat Lab**

# **Sparse matrix factorization for PSFs field estimation**

PSF Meeting, Saclay  
September 27

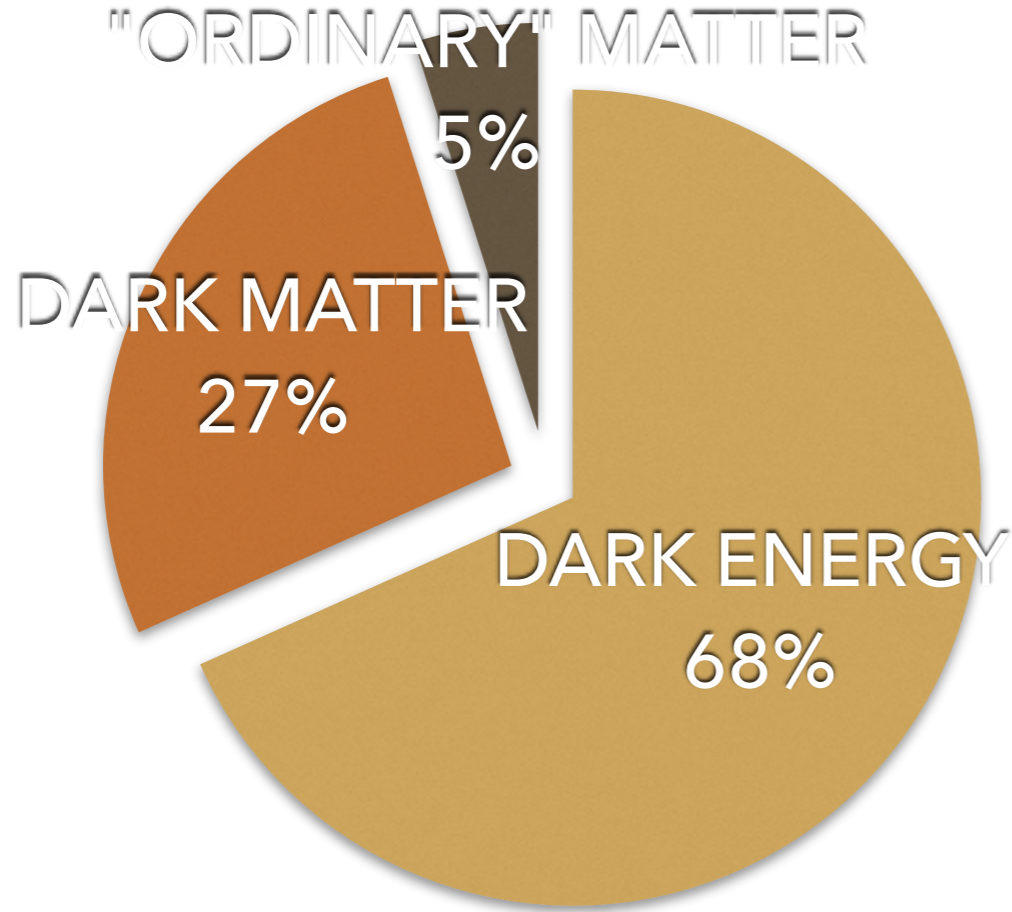
*Fred Ngolè*



# Outline

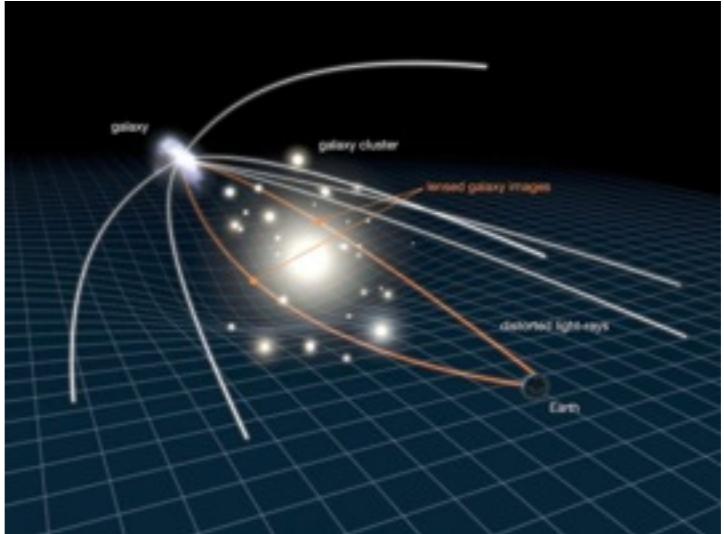
- Context
- Monochromatic PSFs joint super-resolution
- Conclusion

# Context



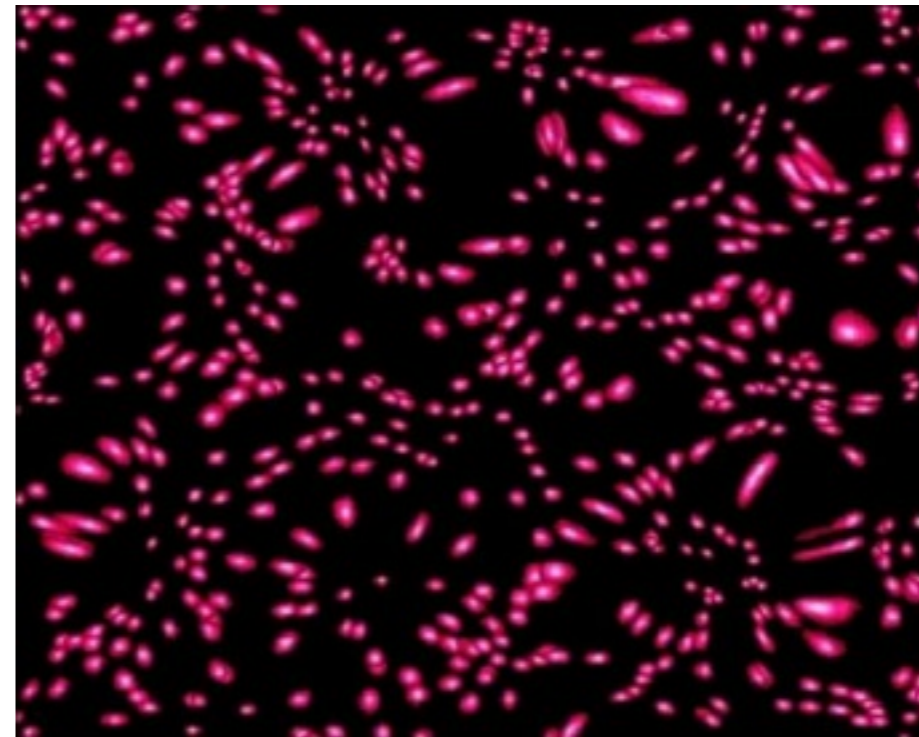
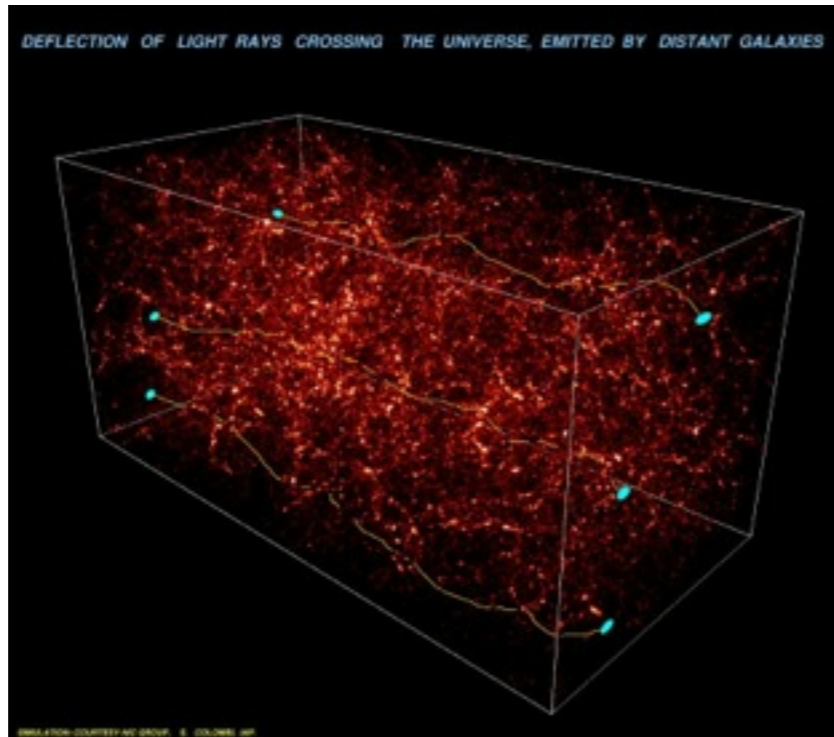
Dark Universe

GR



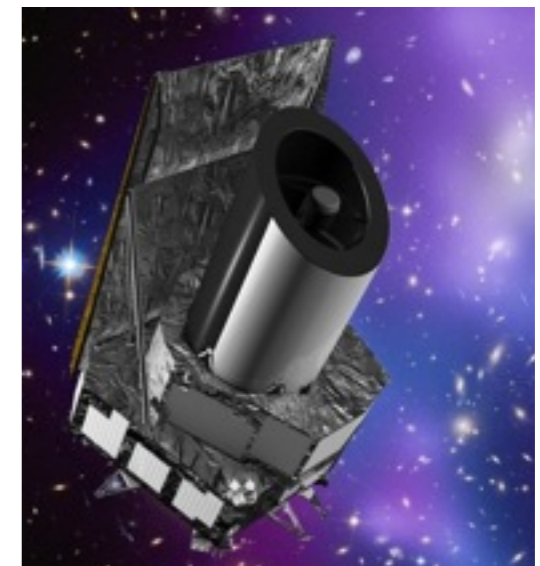
Gravitational lensing

# Context



Weak gravitational lensing  
( $\leq 10\%$  of distortion)

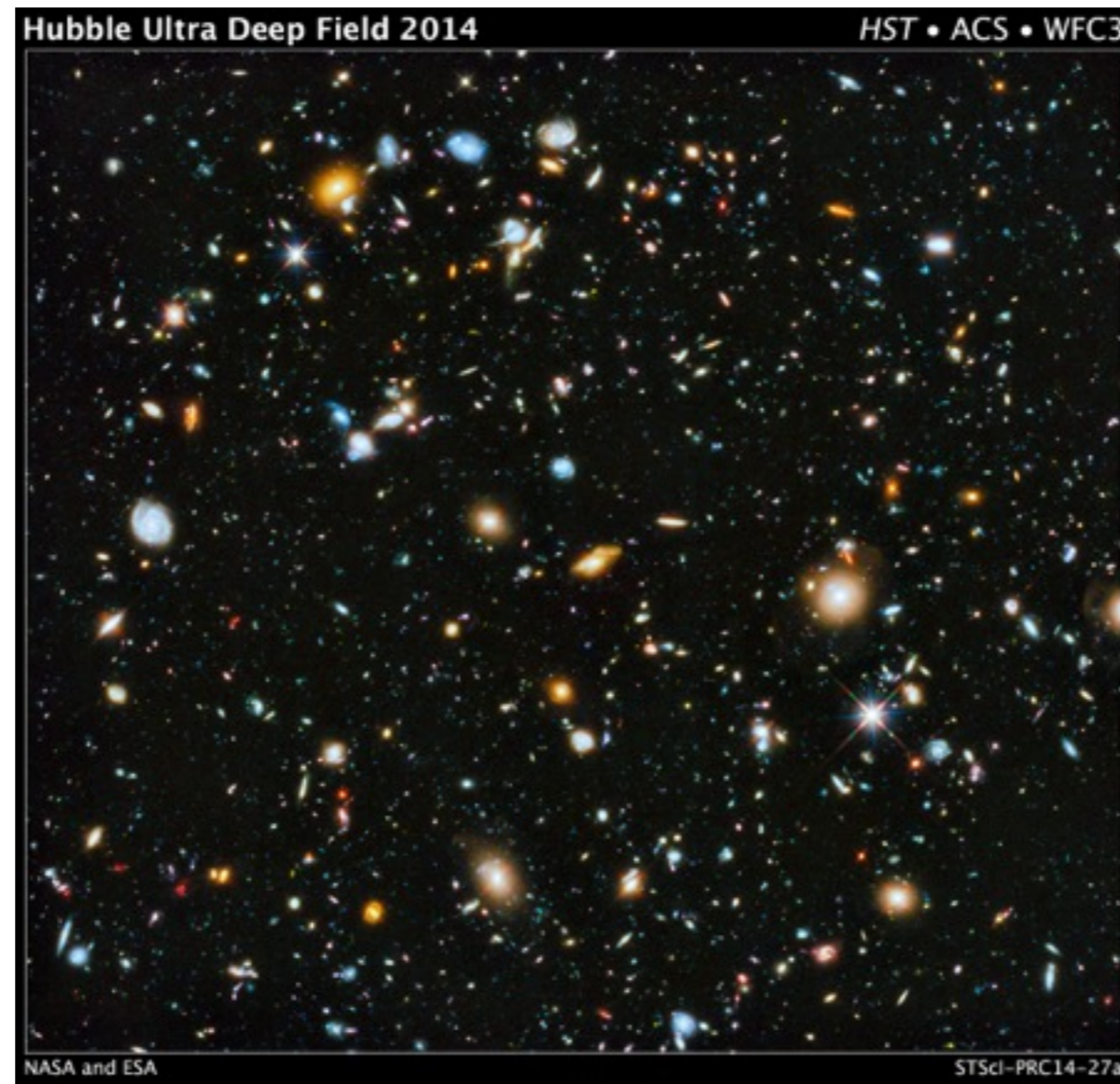
- Data from the ESA Euclid mission  
(to be launched in 2020)





# Context

Weak lensing measurement bottleneck: the PSF



# Context

Model-free PSFs estimation accounting for:

- noise and aliasing
- spatial and temporal variations
- wavelength dependency.

# Monochromatic PSFs joint super-resolution

Observation model

$$\mathbf{y}_k = \mathbf{M}_k \mathbf{x}_k + \mathbf{n}_k, \quad k = 1 \cdots p$$

$\mathbf{y}_k$ :  $k^{\text{th}}$  low resolution image

$\mathbf{M}_k$ : shift and downsampling operator

$\mathbf{x}_k$ :  $k^{\text{th}}$  well resolved image

$\mathbf{n}_k$ : gaussian noise

# Monochromatic PSFs joint super-resolution

## Constraints

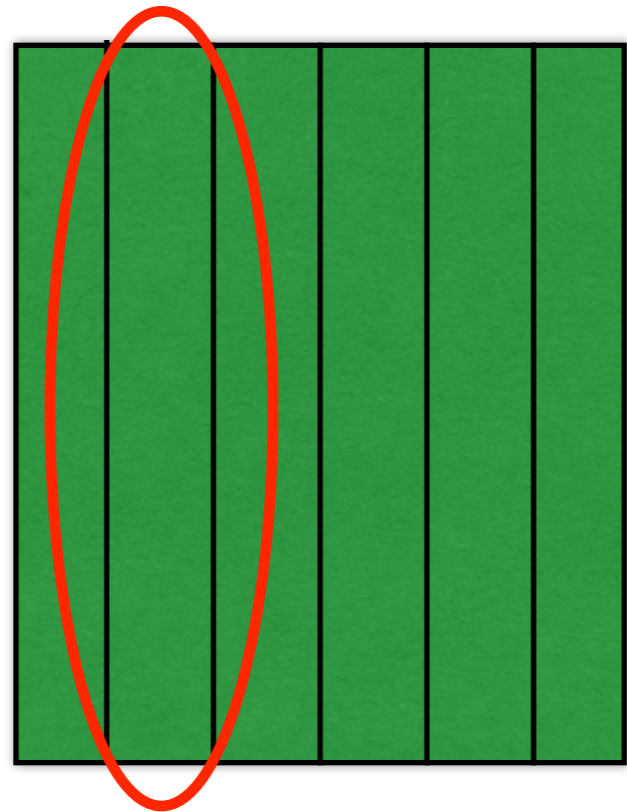
- Piece-wise smoothness
- Low dimensionality
- Spatial regularity
- Positivity



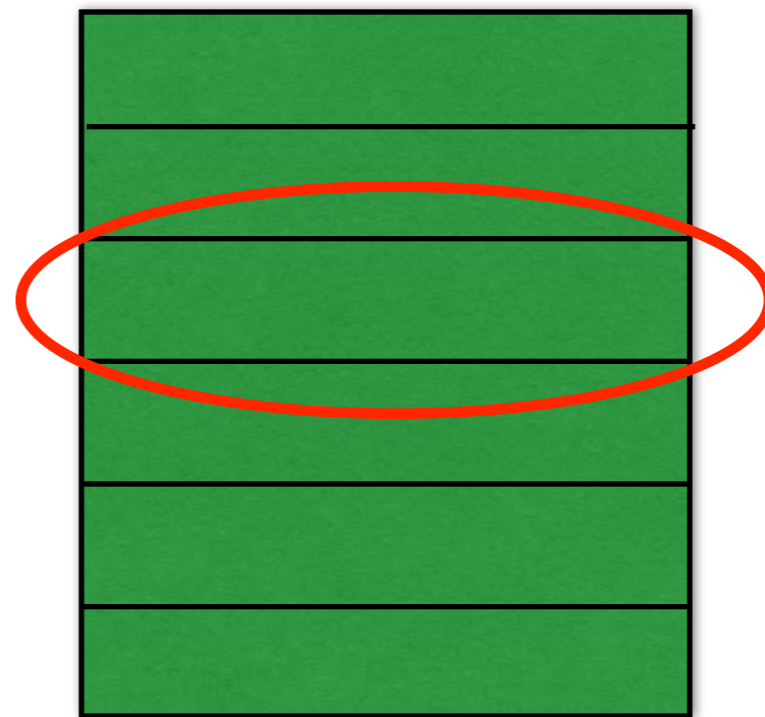
# Monochromatic PSFs joint super-resolution

## Constraints

$$\mathbf{X} = [\mathbf{x}_1, \dots, \mathbf{x}_p] =$$



=



$\Phi \mathbf{X}$  columns sparse

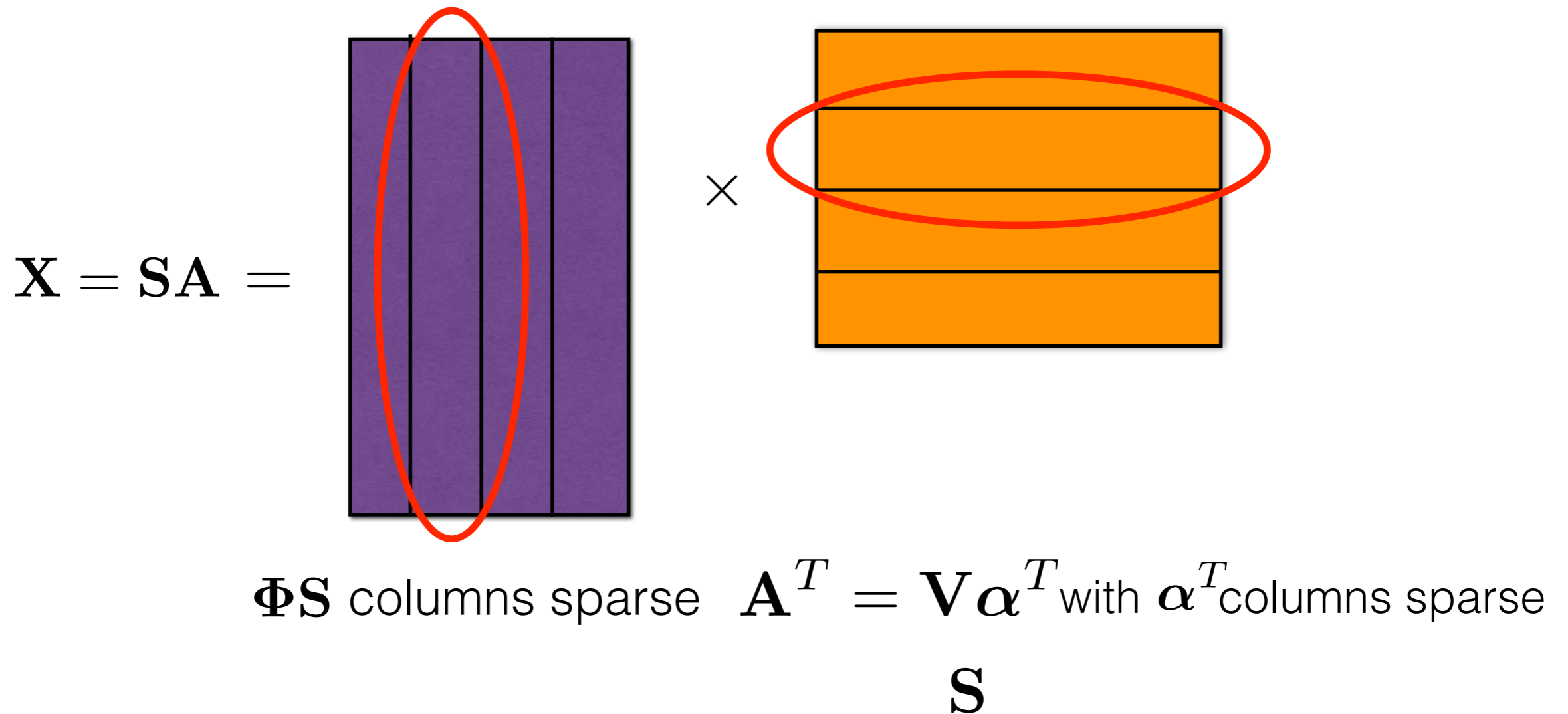
$\mathbf{X}^T = \mathbf{V} \mathbf{W}$  with  $\mathbf{W}$  columns sparse

Pixel domain features dictionary

Spatial frequencies dictionary

# Monochromatic PSFs joint super-resolution

Constraints



# Monochromatic PSFs joint super-resolution

## Spatial frequency dictionary

Observed stars spatial locations:  $(\mathbf{u}_i)_{1 \leq i \leq p}$

*1D case, regular spacing,  $p=2k+1$*

$$\psi_i = \psi_{p-i+1} = -1/|\mathbf{u}_i - \mathbf{u}_{k+1}|^e \quad a > 0, e > 0$$

$$\psi_i = \sum_{\substack{j=1 \\ j \neq k+1}}^p a/|\mathbf{u}_j - \mathbf{u}_{k+1}|^e$$

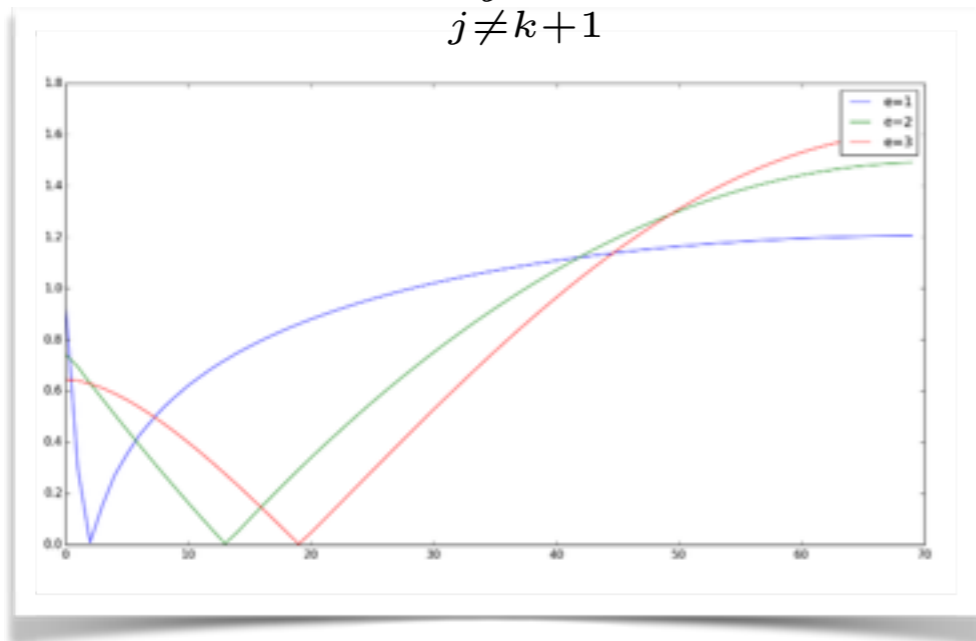
“Notch” filters: varying notch frequency

depending on  $e$  and  $a$

$$\Psi_{e,a}(\mathbf{v}) = \|\mathbf{v} \star \psi_{e,a}\|_2^2$$

$$\Psi : M_{rp}(\mathbb{R}) \mapsto \mathbb{R}^+ , \mathbf{A} \rightarrow \sum_{i=1}^r \Psi_{e_i,a}(\mathbf{A}[i,:])$$

$$0 < e_1 < e_2 < \dots < e_r$$



# Monochromatic PSFs joint super-resolution

Spatial frequency dictionary

$$\widehat{\Psi}_{e,a} : \mathbb{R}^p \mapsto \mathbb{R}^+ , \mathbf{v} \rightarrow \sum_{k=1}^p \left( \sum_{\substack{i=1 \\ i \neq k}}^p \frac{a v_k - v_i}{\|\mathbf{u}_k - \mathbf{u}_i\|_2^e} \right)^2$$

$$\widehat{\Psi}_{e,a}(\mathbf{v}) = \|\mathbf{P}_{e,a} \mathbf{v}\|_2^2$$

$$\mathbf{P}_{e,a}[i,j] = -\frac{1}{\|\mathbf{u}_i - \mathbf{u}_j\|_2^e} \text{ if } i \neq j \quad \mathbf{P}_{e,a}[i,i] = \sum_{\substack{j=1 \\ j \neq i}}^p \frac{a}{\|\mathbf{u}_i - \mathbf{u}_j\|_2^e}$$

$$\widehat{\Psi}_{e,a}(\mathbf{v}) = \mathbf{v}^T \mathbf{Q}_{e,a} \mathbf{v}, \quad \mathbf{Q}_{e,a} = \mathbf{P}_{e,a}^T \mathbf{P}_{e,a} = \mathbf{V}_{e,a} \mathbf{D}_{e,a} \mathbf{V}_{e,a}^T$$

$$\Psi(\mathbf{A}) = \sum_{i=1}^r \sum_{j=1}^p \mathbf{d}_{e_i,a}[j] \langle \mathbf{v}, \mathbf{V}_{e_i,a}[:,j] \rangle^2$$

# Monochromatic PSFs joint super-resolution

Spatial frequency dictionary

$$(e_i, a_i)_{1 \leq i \leq r} \quad \mathbf{V} = [\mathbf{V}_{e_1, a_1}, \dots, \mathbf{V}_{e_r, a_r}]$$

$\mathbf{P}_{e,1}$ : laplacian matrix of a graph

# Monochromatic PSFs joint super-resolution

Optimization problem

$$\Omega_1 = \{\boldsymbol{\alpha} \in M_{r,N}(\mathbb{R}) / \|\boldsymbol{\alpha}[l, :]\|_0 \leq \eta_l, l = 1 \dots r\}$$

$$\Omega_2 = \{(\mathbf{S}, \boldsymbol{\alpha}) \in M_{nr}(\mathbb{R}) \times M_{r,N}(\mathbb{R}) / \mathbf{S}\boldsymbol{\alpha}\mathbf{V}^T \succeq_{M_{np}(\mathbb{R})} \mathbf{0}\}$$

$$\min_{\boldsymbol{\alpha}, \mathbf{S}} \frac{1}{2} \|\mathbf{Y} - \mathcal{F}(\mathbf{S}\boldsymbol{\alpha}\mathbf{V}^T)\|_F^2 + \sum_{i=1}^r \|\mathbf{w}_i \odot \boldsymbol{\Phi}_s \odot \mathbf{s}_i\|_1 + \iota_{\Omega_1}(\boldsymbol{\alpha}) + \iota_{\Omega_2}(\mathbf{S}, \boldsymbol{\alpha})$$



# Monochromatic PSFs joint super-resolution

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## Algorithm 1 Resolved components analysis (RCA)

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- 1: Parameters estimation and initialization:  
 Harmonic constraint parameters  $(e_i, a_i)_{1 \leq i \leq r} \rightarrow \mathbf{V}, \mathbf{A}_0$   
 Noise level,  $\mathbf{A}_0 \rightarrow \mathbf{W}_{0,0}$
  - 2: Alternate minimization
  - 3: **for**  $k = 0$  to  $k_{\max}$  **do**
  - 4:   **for**  $j = 0$  to  $j_{\max}$  **do**
  - 5:      $\mathbf{S}_k = \underset{\mathbf{S}}{\operatorname{argmin}} \frac{1}{2} \|\mathbf{Y} - \mathcal{F}(\mathbf{S}\mathbf{A}_k)\|_F^2 + \sum_{i=1}^r \|\mathbf{W}_{k,j}[:, i] \odot \Phi_s \mathbf{S}[:, i]\|_1$  s.t.  $\mathbf{S}\mathbf{A}_k \geq 0$
  - 6:     update:  $\mathbf{W}_{k,0}, \mathbf{S}_k \rightarrow \operatorname{update}(\mathbf{W}_{k,j+1})$
  - 7:   **end for**
  - 8:    $\boldsymbol{\alpha}_{k+1} = \underset{\boldsymbol{\alpha}}{\operatorname{argmin}} \frac{1}{2} \|\mathbf{Y} - \mathcal{F}(\mathbf{S}_k \boldsymbol{\alpha} \mathbf{V}^T)\|_F^2$  s.t.  $\|\boldsymbol{\alpha}[l, :]\|_0 \leq \eta_l$
  - 9:   update: Noise level,  $\boldsymbol{\alpha}_{k+1} \rightarrow \mathbf{W}_{k+1,0}$
  - 10:    $\mathbf{A}_{k+1} = \boldsymbol{\alpha}_{k+1} \mathbf{V}^T$
  - 11:    $\mathbf{A}_{k+1}[i, :] = \mathbf{A}_{k+1}[i, :] / \|\mathbf{A}_{k+1}[i, :]\|_2$ , for  $i = 1 \dots r$
  - 12: **end for**
  - 13: **Return:**  $\mathbf{S}_{k_{\max}}, \mathbf{A}_{k_{\max}}$ .
-

# Monochromatic PSFs joint super-resolution

## Numerical results

### *Data*

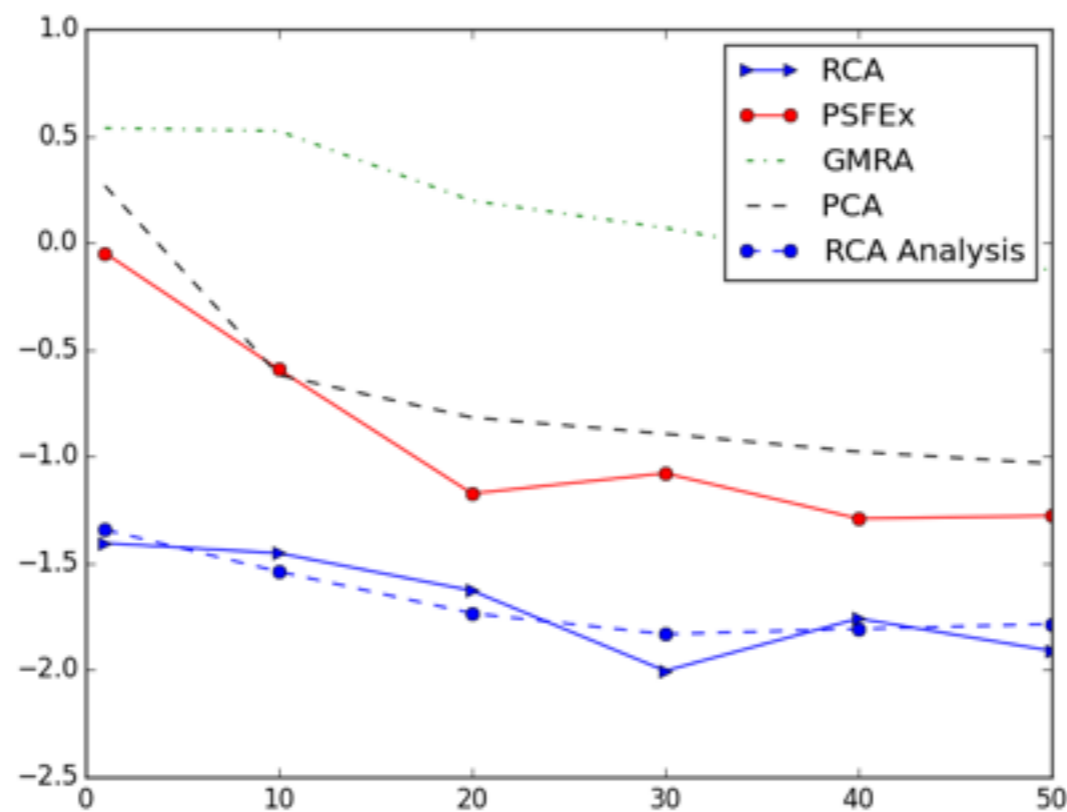
- 500 42x42 simulated Euclid telescope optical PSFs
- Realistic effects included such polishing and alignment defects
- Upsampling factor of 2 in the SR case

# Monochromatic PSFs joint super-resolution

## Numerical results

*Dimension reduction*

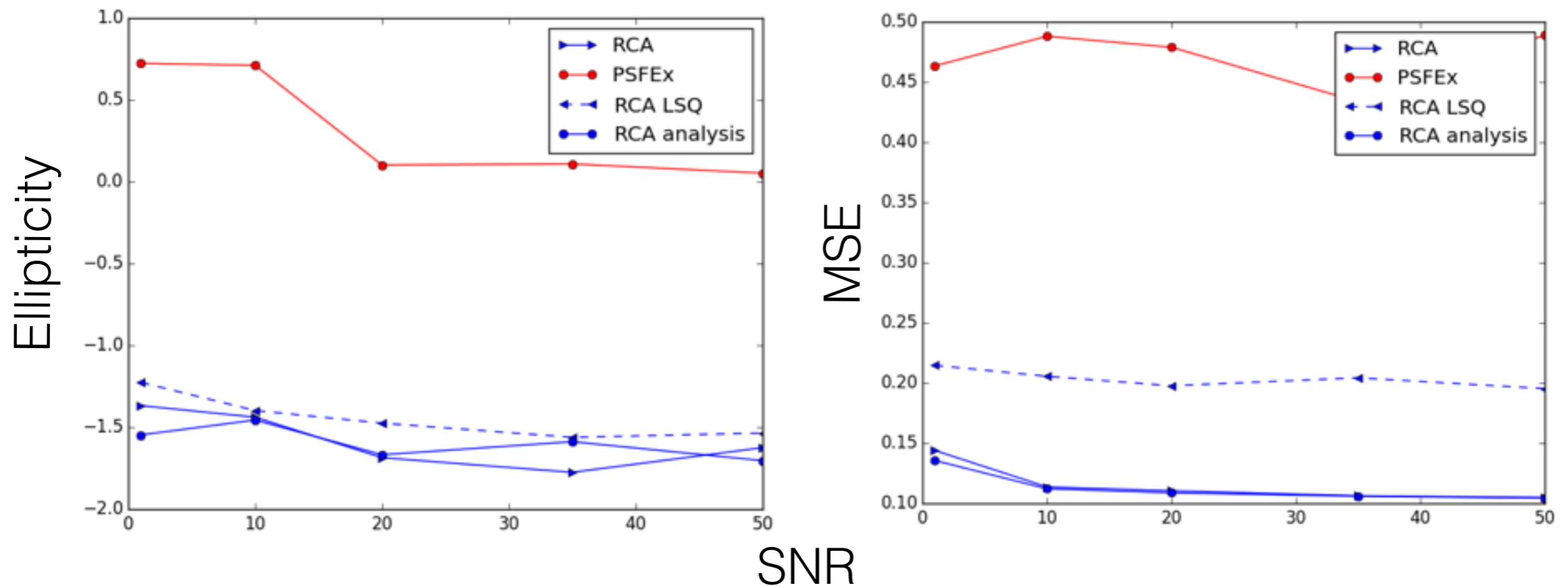
Ellipticity  
vector error



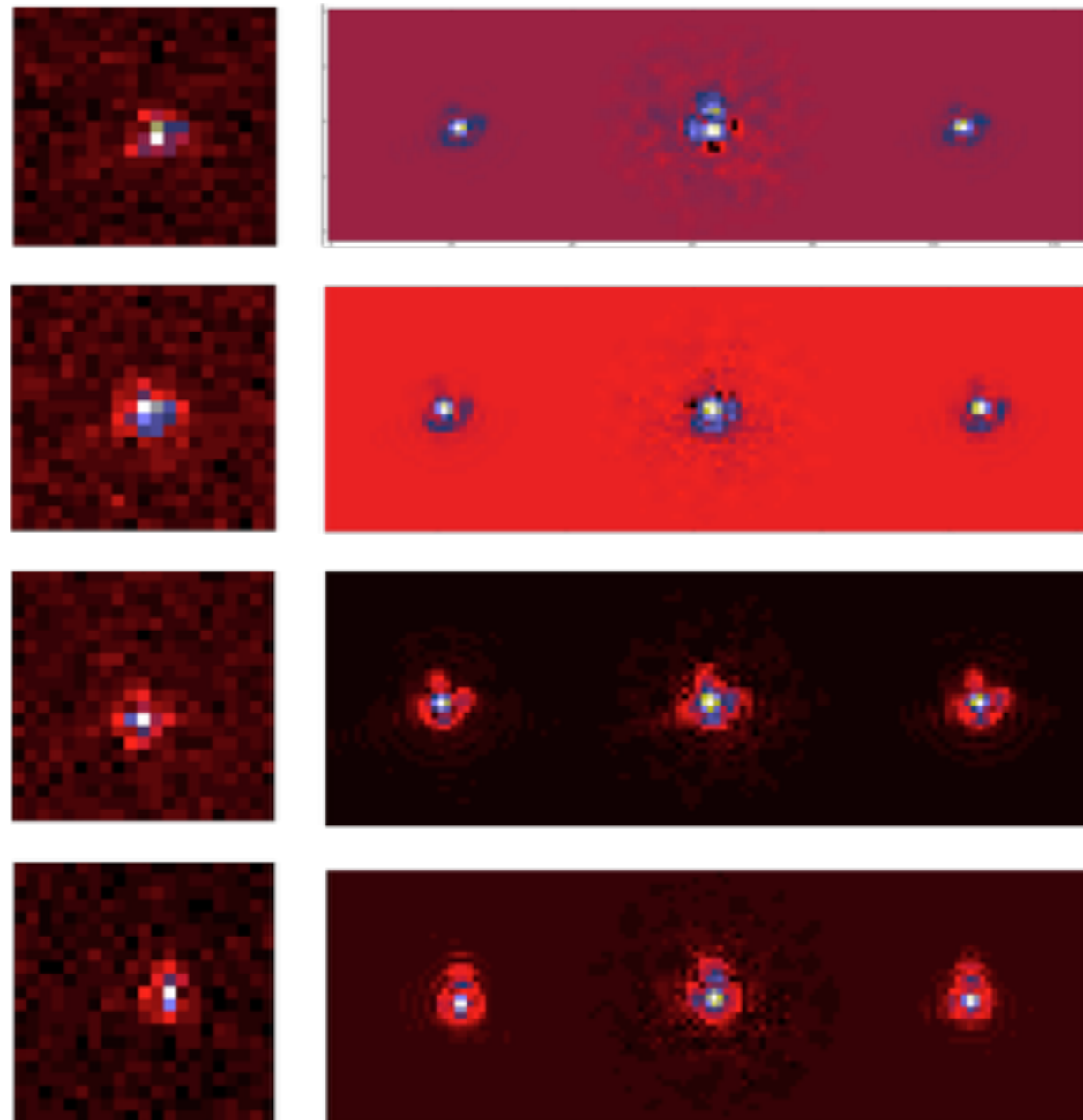
# Monochromatic PSFs joint super-resolution

## Numerical results

*Dimension reduction and super-resolution*



# Monochromatic PSFs joint super-resolution



# Conclusion

- Noise robust dimension reduction and super-resolution method taking advantage of the PSFs field spatial regularity
- Good accuracy on both PSFs shape and pixels values



# Bibliography

- F. M. Ngolè Mboula, J.-L. Starck, K. Okomura, J. Amiaux, P. Hudelot. "*Constraint matrix factorization for space variant PSFs field restoration*", Inverse Problems, In press.
- Condat, Laurent. "*A primal–dual splitting method for convex optimization involving Lipschitzian, proximable and linear composite terms.*" Journal of Optimization Theory and Applications 158.2 (2013): 460-479..

*Thank you for your attention!*